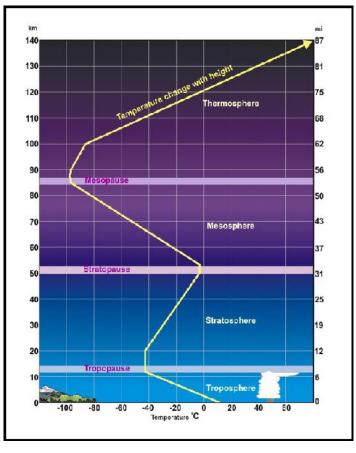
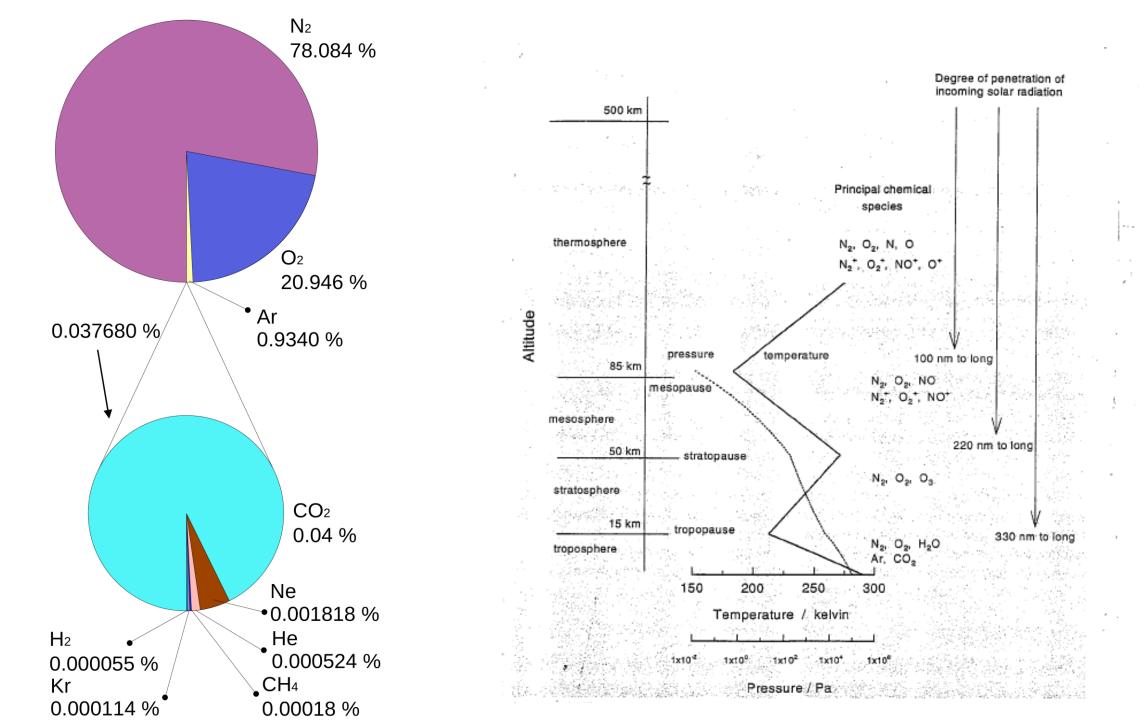
Structure and Composition of Earth's Atmosphere









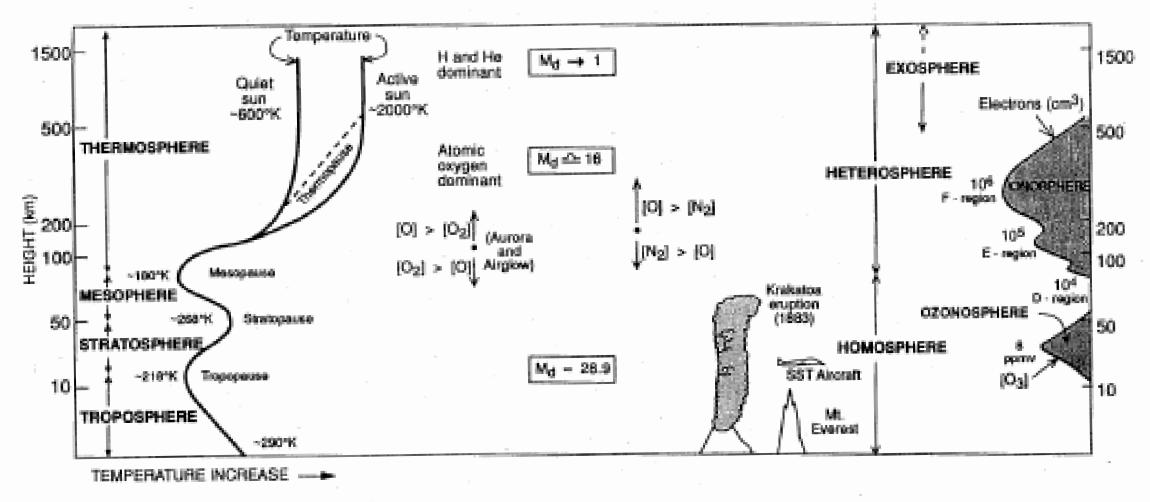


Figure 3.1. Two ways of dividing the atmosphere: by temperature structure (left side) and by composition (right side). The change in the apparent molecular weight of air (M_d) due to the changing composition of the atmosphere with height is shown in the center.

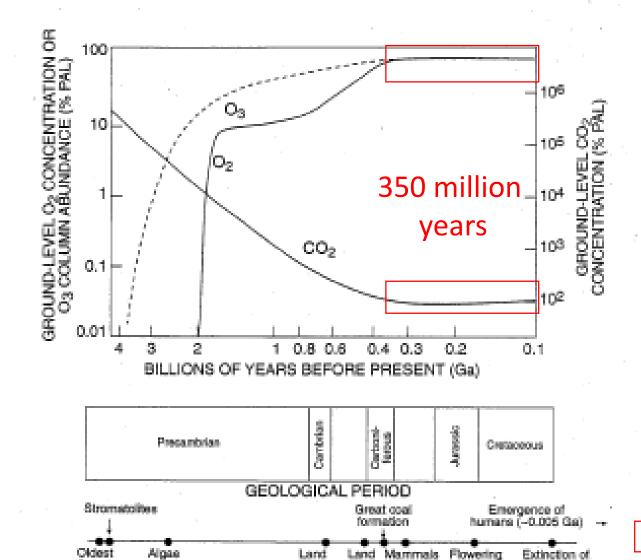


Figure 1.2. Schematic diagram showing predictions of the evolution of oxygen, ozone and carbon dioxide to present atmospheric levels (PAL). [After R. P. Wayne, Chemistry of Atmospheres, Oxford University Press, p. 404 (1991) by permission of Oxford University Press; and, J. F. Kasting, personal communication (1999).]

animale

plants

dinosaurs

(-0.068 Ge)

sedimentary

rocks

Biological Processes

$$H_2O + CO_2 \rightarrow O_2 + \{CH_2O\}$$

hv (sunlight)

Respiration

Combustion

Table 1.1. Composition of dry unpolluted air by volume

Miles Says and Says	
Nitrogen	78.084%
Oxygen	20.946%
Argon	0.934%
Carbon dioxide	360 ppm (variable)
Neon	18.18 ppm
Helium	5.24 ppm
Methane	1.6 ppm
Krypton	1.14 ppm
Hydrogen	0.5 ppm
Nitrous Oxide	0.3 ppm
Xenon	0.087 ppm
The second secon	

Table 1.4. Size and general mixing of various reservoirs. Plants, animals and organic matter are included in the biosphere, but not coal or sedimentary carbon.

ē.,	mass (kg)	mixing time (a)
Biosphere	4.2 × 1015	60
Atmosphere	5.2×10^{18}	0.2 ~ 60-80 days
Hydrosphere	1.4 × 10 ²¹	1600
Crust	2.4×10^{32}	$> 3 \times 10^7$
Mantle	4.0×10^{24}	> 108
Core	1.9×10^{24}	

Table 1.5. Standard properties of the atmosphere at sea level

	Density	1.2250014 kg m ⁻³
	Gravitational acceleration (g)	9.80665 m s ⁻²
	Kinematic viscosity	1.4607 × 10 ⁻⁵ m ² s ⁻¹
66 nm → Mean free pa	Mean free path	$6.632 \times 10^{-8} \mathrm{m}$
	Molecular weight (M _m)	28.966
	Number density (n)	$2.5476 \times 10^{19} \text{cm}^{-3}$
	Pressure (p)	101325 Pa
	Scale height (H)	8434 m
0.37 P°		

 $P_H = P^o e^{-1} = 1/e P^o = 0.37 P^o$

Table 2.1. Residence times of some atmospheric gases^e (in many cases only very rough estimates are possible)

Gas	Residence Time
Nitrogen (N ₂)	1.6×10^{7} a
Helium (He)	10 ⁶ a
Oxygen (O ₂)	3,000-10,000 a
Carbon dioxide (CO ₂)	3-4 a
Nitrous oxide (N ₂ O)	150 a
Methane (CH ₄)	9 a
CFC-12 (CF ₂ Cl ₂)	>80a
CFC-11 (CFCI ₃)	-80 a
Hydrogen (H₂)	4-8 a
Methyl chloride (CH ₃ Cl)	2-3 a
Carbonyl sulfide (COS)	~2 a
Ozone (O ₃)	100 days
Carbon disulfide (CS ₂)	40 days
Carbon monoxide (CO)	~60 days
Water vapor	~10 days
Formaldehyde (CH ₂ O)	5-10 days
Sulfur dioxide (SO ₂)	1 day
Ammonia + Ammonium (NH ₃ + NH ₂)	2-10 days
Nitrogen dioxide (NO ₂)	0.5-2 days
Nitrogen oxide (NO)	0.5-2 days
Hydrogen chloride (HCl)	4 days
Hydrogen sulfide (H-S)	1-5 days
Hydrogen peroxide (H ₂ O ₂)	1 day
Dimethyl sulfide (CH ₂ SCH ₃)	0.7 days

[&]quot;The residence time (or lifetime) is defined as the amount of the chemical in the atmosphere divided by the rate at which the chemical is removed from the atmosphere. This time scale characterizes the rate of adjustment of the atmospheric concentration of the chemical if the emission rate is changed suddenly.

Chemically 'inert'

Chemically 'reactive'

^{*}The residence time of liquid water in clouds is ~6 h.

Earth's atmosphere temperature (°F) -140-120-100-80 -60 -40 -20 Percentages at selected -70 altitudes are 110percentages -65 of sea-level 100pressure. -60 thermosphere ionosphere. 90 km, 0.0001% magnetosphere 90begin -55mesopause -5080--45 -45 (sejilli) altitude (km) mesosphere 50 km, 0.1% -35 stratopause -30 40--25 -2030 stratosphere ozone layer (main concentration) -15Mount Everest 20-8.85 km, 28% -10tropopause troposphere -100-90-80-70-60-50-40-30-20-10 250 500 750 1,000 20 temperature (°C) pressure (millibars) © 2007 Encyclopædia Britannica, Inc.

Thermal Structure

Radiative Energy Balance
Energy in (absorbed from Sun) =
Energy out (emitted by Earth)

Blackbody radiation depends only on temperature of emitter

Sun ~ 6000 K Earth ~ 255 K

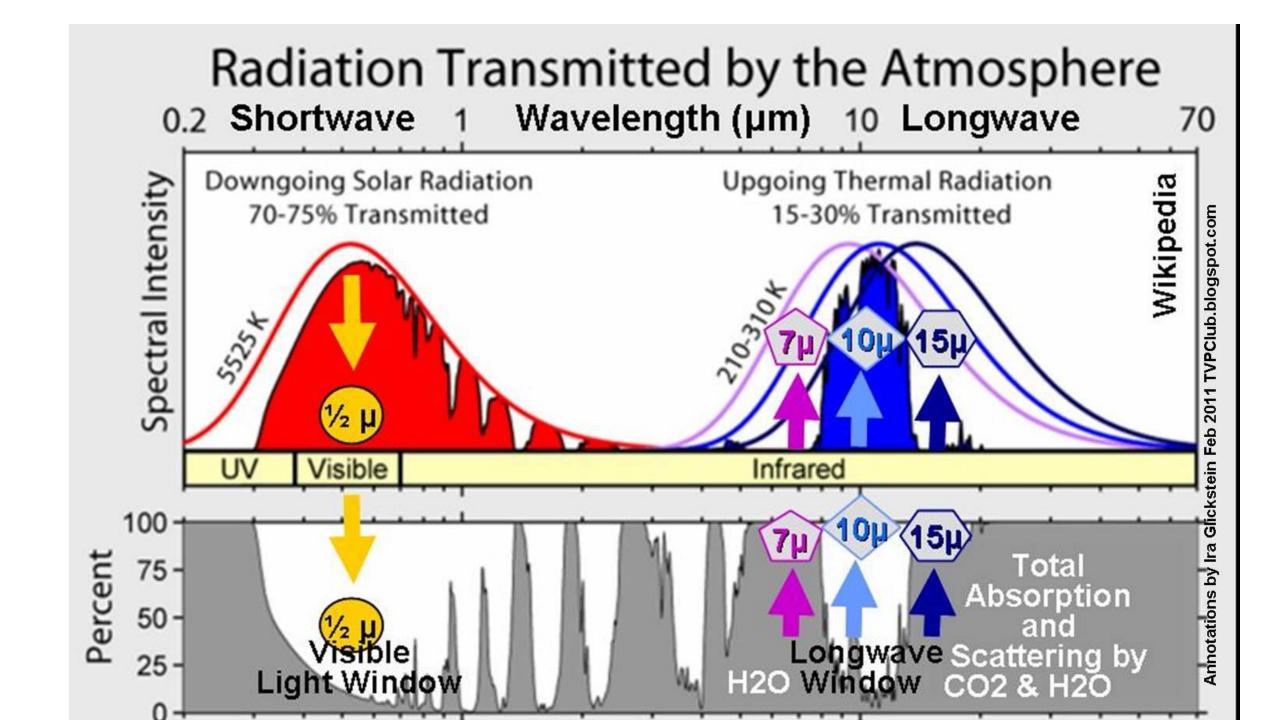


Fig. 1.2. A comparison between (a) the electromagnetic spectrum for black bodies at 6000 and 255 K and (b) the absorption spectrum of gases in the Earth's atmosphere. Note that the atmosphere is practically transparent to black body radiation emitted at temperatures typical of the sun.

